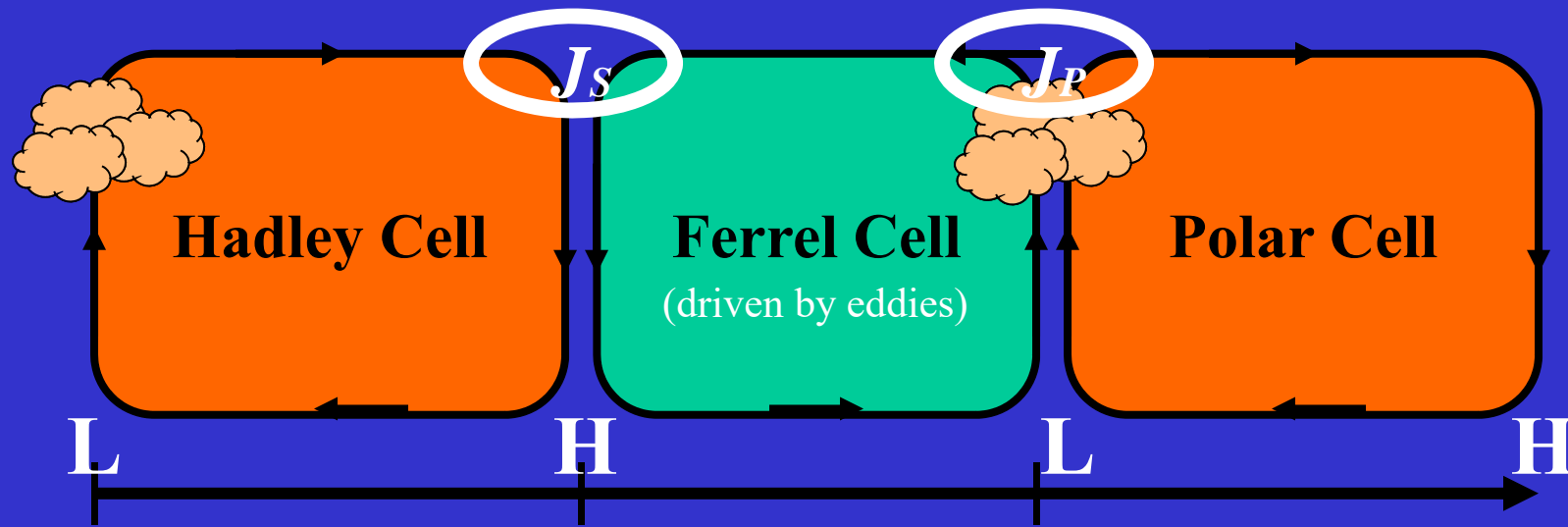


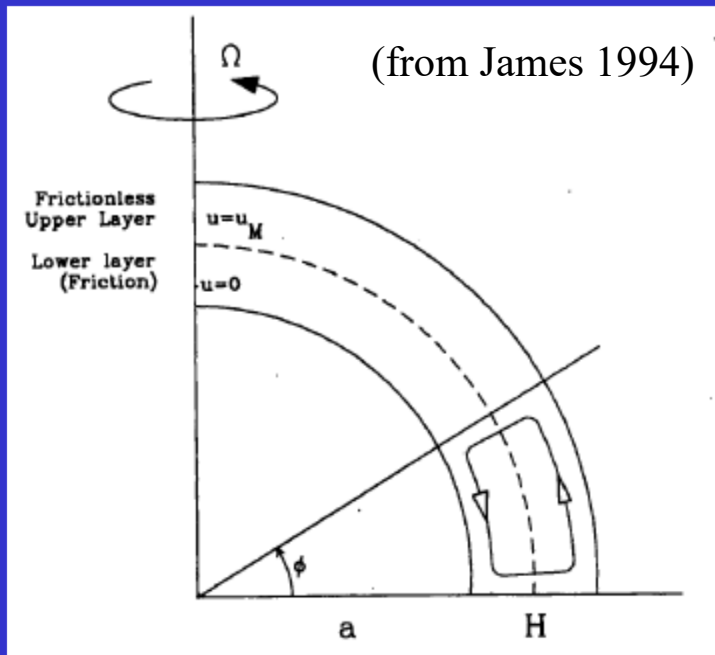
Lecture 11: Atmospheric General Circulation



- Hadley Cell – a thermal-driven circulation
- Farrell Cell – an eddy-driven circulation



The Held-Hou Model (1980) for Hadley Circulation



thermal wind balance

$$\frac{\partial U}{\partial z} \propto \frac{\partial T}{\partial y}$$

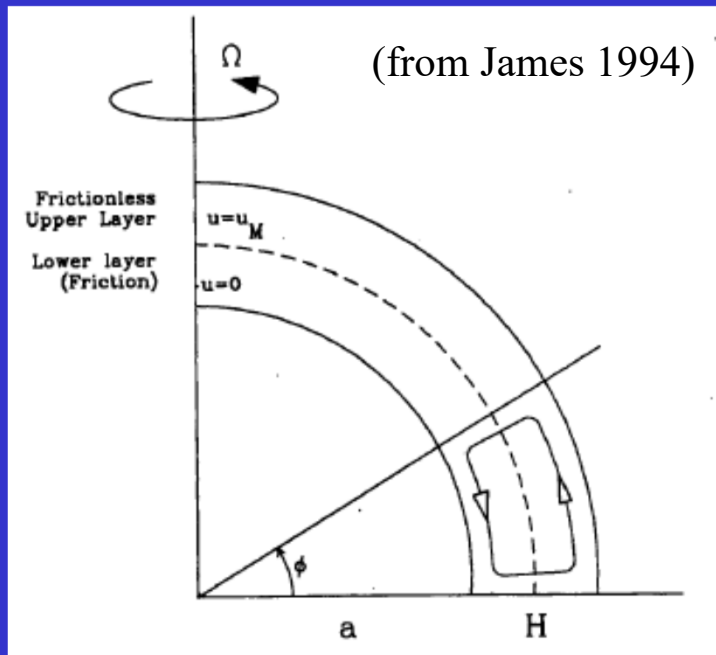
angular momentum conservation

radiative equilibrium

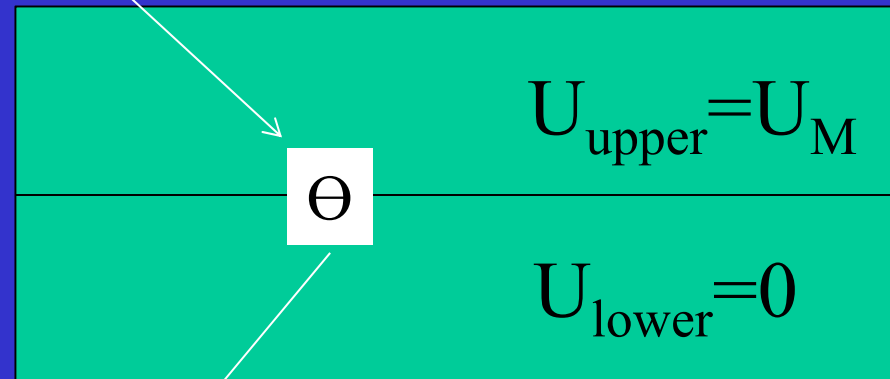
- The Held-Hou model is a two-level model on the sphere with equatorward flow at the surface and poleward flow at height H .
- The model uses angular momentum conservation and thermal wind balance to determine the width and strength of the Hadley circulation.



Held and Hou Model: Radiative Forcing



Potential tempter at middle-level



$$\frac{D\theta}{Dt} = \frac{\theta_E - \theta}{\tau_E},$$

$$\theta_E(\varphi) = \theta_0 - \frac{1}{3} \Delta\theta (3 \sin^2 \varphi - 1).$$

$$\varphi \approx \sin \varphi \approx y/a$$

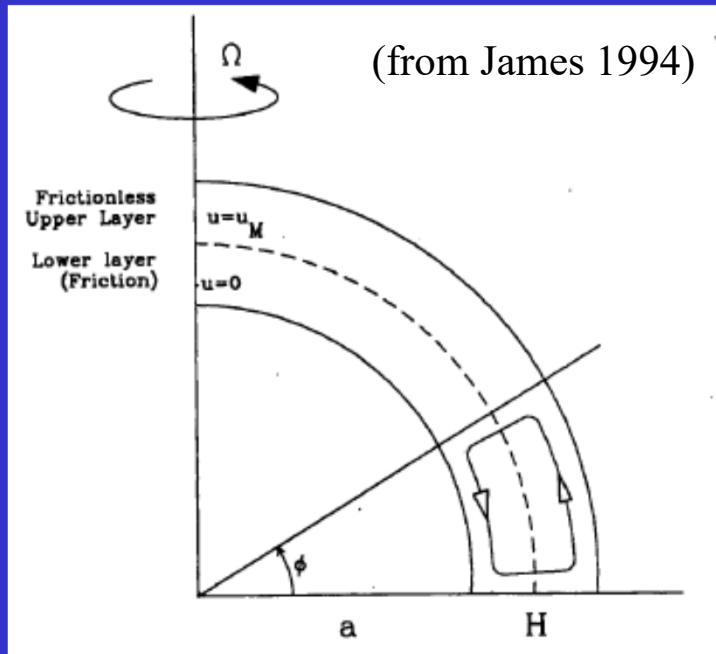
$$\theta_E(y) = \theta_{E0} - \Delta\theta \frac{y^2}{a^2},$$

where $\theta_{E0} = \theta_0 + \Delta\theta/3$

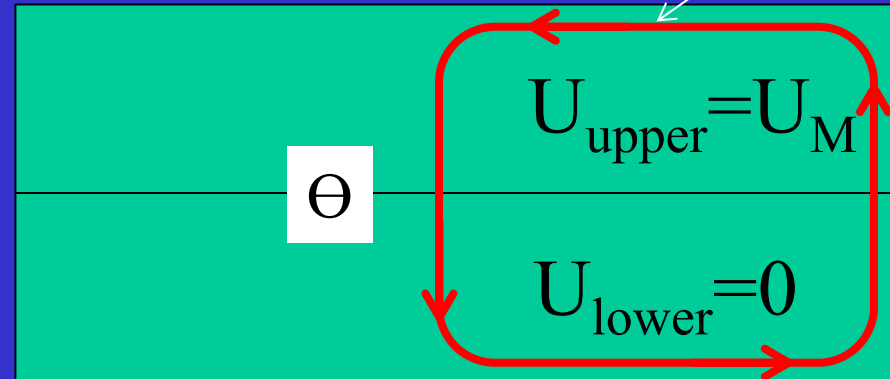
θ_E : global mean radiative equilibrium temperature
 $\Delta\theta$: pole-to-equator temperature deference

- Radiative processes are represented in the model using a Newtonian cooling formulation.

Held and Hou Model: Dynamics



Hadley Circulation



$(\Omega a \cos \varphi + u)a \cos \varphi$ absolute angular momentum

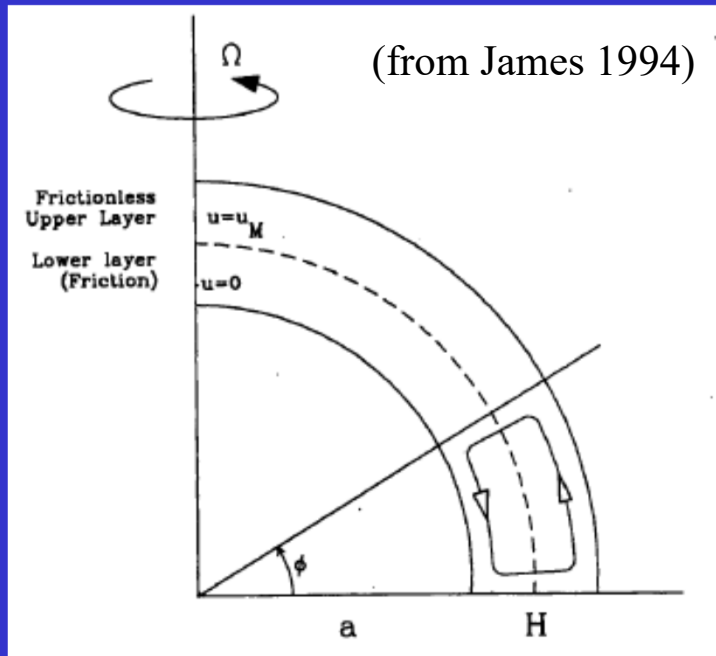
$u(\varphi) = \Omega a \frac{\sin^2 \varphi}{\cos \varphi}$ angular momentum conservation from equator to a latitude φ

↓ $\varphi \approx \sin \varphi \approx y/a$

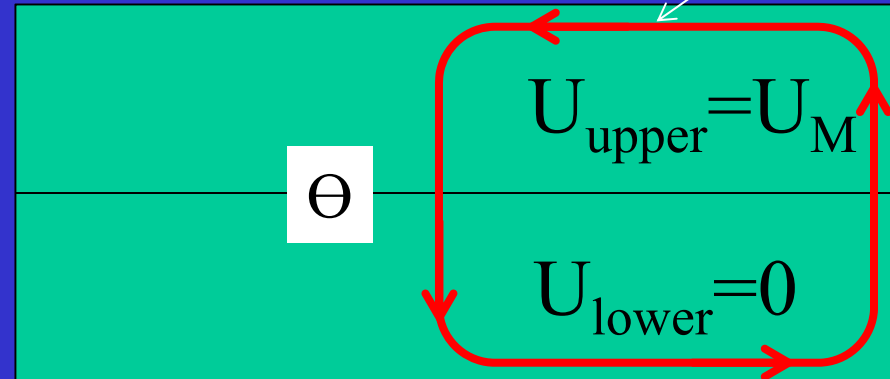
$$U_M = \frac{\Omega}{a} y^2.$$

- The zonal-mean zonal wind (U) is determined by the conservation of absolute angular momentum.

Held and Hou Model: Thermal Wind Balance



Hadley Circulation



$$\frac{\partial u}{\partial z} = \frac{U_M}{H} = \frac{\Omega}{aH} y^2.$$

$$\frac{\partial \theta}{\partial y} = -\frac{2\Omega^2 \theta_0}{a^2 g H} y^3.$$

$$\theta_M = \theta_{M0} - \frac{\Omega^2 \theta_0}{2a^2 g H} y^4,$$

based on thermal wind balance

$$\frac{\partial \mathbf{V}_g}{\partial \ln p} = -\frac{R}{f} \mathbf{k} \times \nabla_p T$$

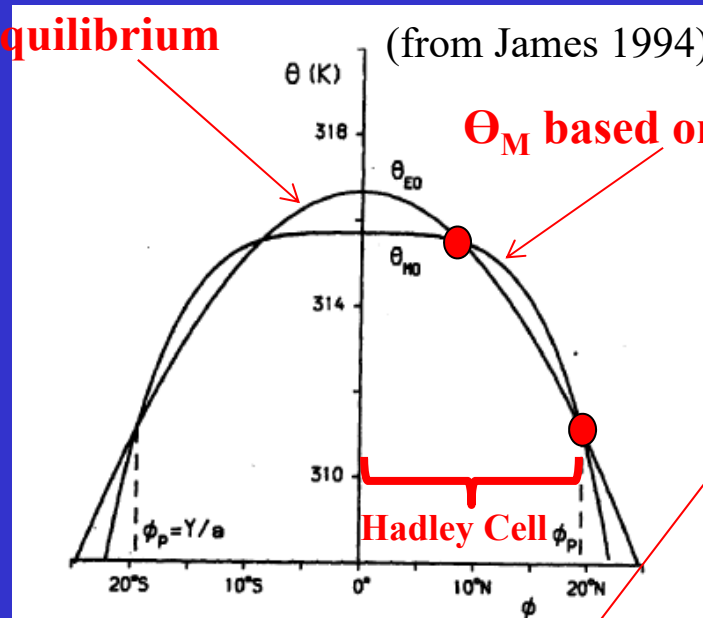
Θ_M : potential temperature derived based on conservation of angular momentum

- The potential temperature structure at the middle level of the model is calculated from thermal wind balance.

Held and Hou Model: Hadley Circulation

Θ_E based on radiative equilibrium

$$\theta_E(y) = \theta_{E0} - \Delta\theta \frac{y^2}{a^2},$$



Θ_M based on momentum conservation

$$\theta_M = \theta_{M0} - \frac{\Omega^2 \theta_0}{2a^2 g H} y^4,$$

$$\frac{D\theta}{Dt} = \frac{\theta_E - \theta}{\tau_E},$$

- The steady state solution to the above equation can be found from the cross points of Θ_E and Θ_M .
- Zone 1 (Equator to the first crossing point): $\Theta_E > \Theta_M \rightarrow$ radiative heating
- Zone 2 (first to second crossing point): $\Theta_E < \Theta_M \rightarrow$ radiative cooling
- Hadley cell is driven by radiative heating to move from equator to Y while conserving angular momentum.

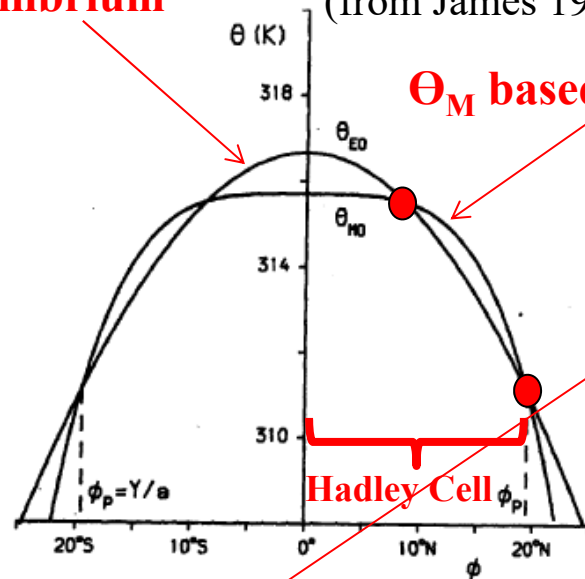
Held and Hou Model: HC Width Y

Θ_E based on radiative equilibrium

$$\theta_E(y) = \theta_{E0} - \Delta\theta \frac{y^2}{a^2},$$

$$\theta_{E0} = \theta_0 + \Delta\theta/3$$

(from James 1994)



Θ_M based on momentum conservation

$$\theta_M = \theta_{M0} - \frac{\Omega^2 \theta_0}{2a^2 g H} y^4,$$

Assume continuity of potential temperature at $y = Y$
 $\rightarrow \Theta_E(Y) = \Theta_M(Y)$

$$\frac{D\theta}{Dt} = \frac{\theta_E - \theta}{\tau_E},$$

$$\int_0^Y \frac{D\theta}{Dt} dy = 0,$$

Since the model assumes a steady state, there can be no net heating of an air parcel when it completes a circuit of the Hadley cell

$$\int_0^Y \theta_M dy = \int_0^Y \theta_E dy.$$

$$\theta_{M0} - \frac{\Omega^2 \theta_0}{2a^2 g H} Y^4 = \theta_{E0} - \frac{\Delta\theta}{a^2} Y^2.$$

$$\theta_{M0} - \frac{\Omega^2 \theta_0}{10a^2 g H} Y^4 = \theta_{E0} - \frac{\Delta\theta}{3a^2} Y^2.$$

$$Y = \left(\frac{5\Delta\theta g H}{3\Omega^2 \theta_0} \right)^{1/2},$$

HC Width

$$\theta_{M0} = \theta_{E0} - \frac{5\Delta\theta^2 g H}{18a^2 \Omega^2 \theta_0}.$$

HC Strength

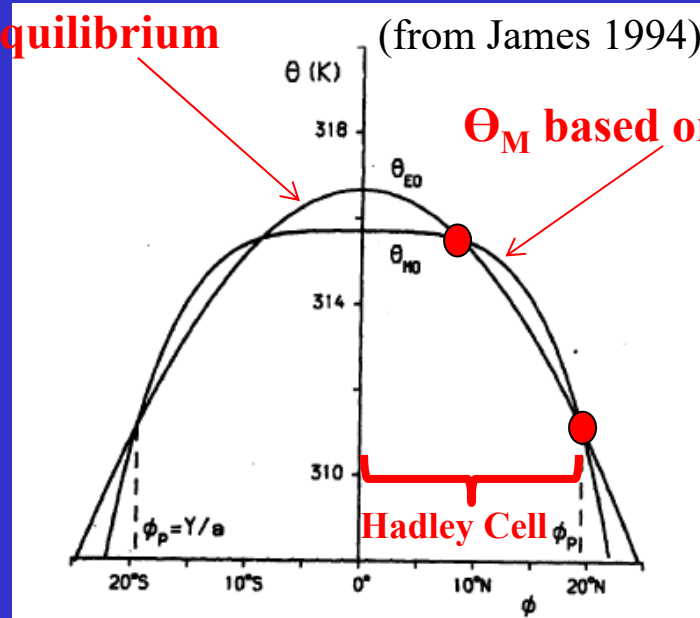
Held and Hou Model: Hadley Circulation

Θ_E based on radiative equilibrium

$$\theta_E(y) = \theta_{E0} - \Delta\theta \frac{y^2}{a^2},$$

$$\theta_{E0} = \theta_0 + \Delta\theta/3$$

(from James 1994)



Θ_M based on momentum conservation

$$\theta_M = \theta_{M0} - \frac{\Omega^2 \theta_0}{2a^2 g H} y^4,$$

$$Y = \left(\frac{5\Delta\theta g H}{3\Omega^2 \theta_0} \right)^{1/2},$$

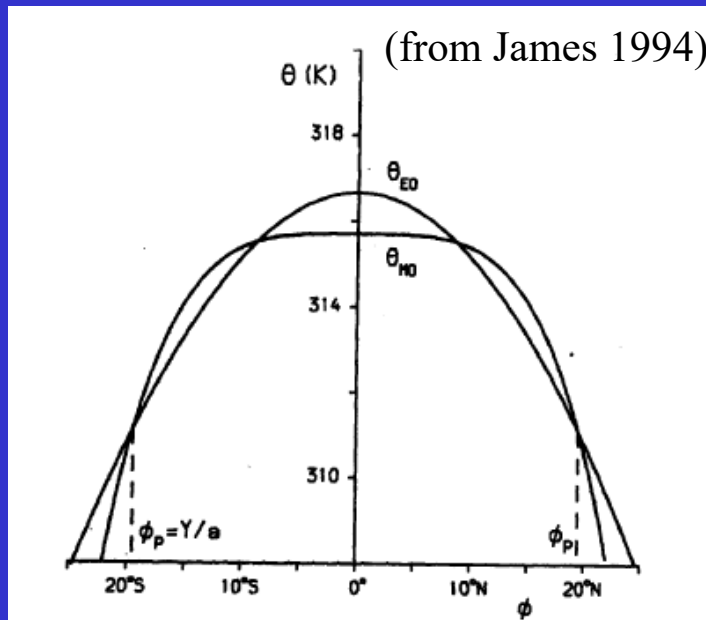
HC Width \leftarrow Come close to the observations (too narrow)

$$\theta_{M0} = \theta_{E0} - \frac{5\Delta\theta^2 g H}{18a^2 \Omega^2 \theta_0}.$$

HC Strength \leftarrow Too weak!



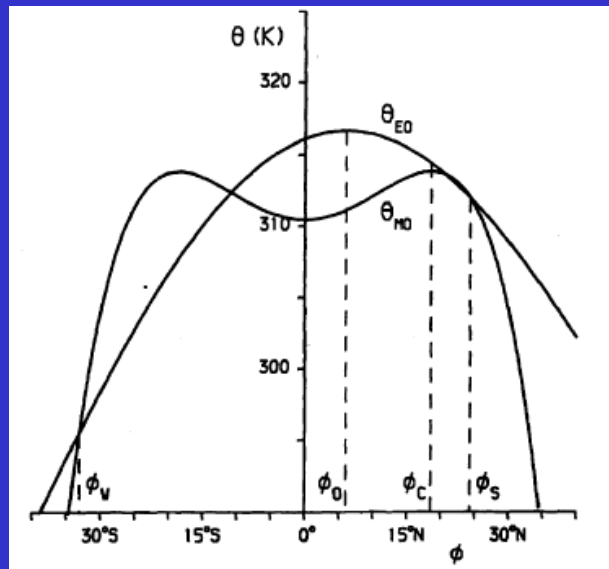
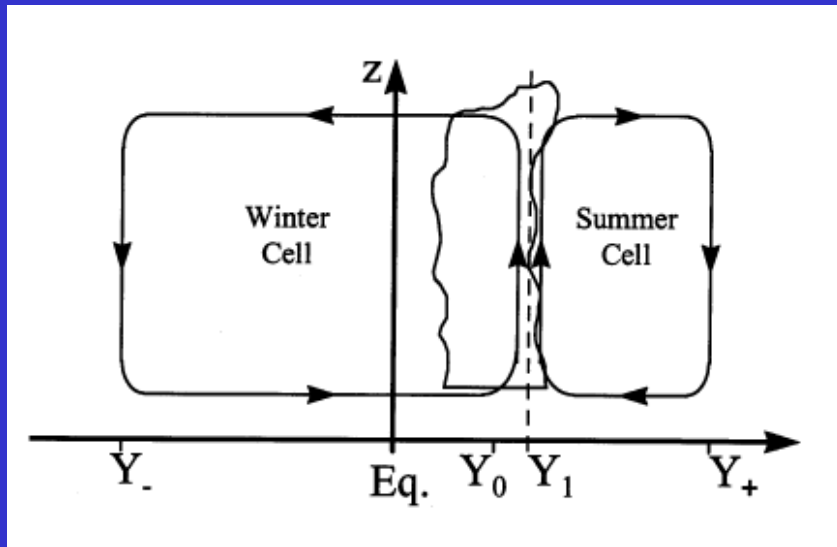
Factors Affect the Hadley Circulation Width



$$Y = \left(\frac{5\Delta\theta g H}{3\Omega^2 \theta_0} \right)^{1/2},$$

- ❑ The Held-Hou model predicts that the width of the Hadley cell is inversely proportional to the planetary rotation rate.
- ❑ At low rotation rates the Hadley cells extend far poleward and account for most of the heat transport from equator to pole.
- ❑ At high rotation rates the Hadley cells are coned near the equator and baroclinic waves poleward of the Hadley circulations are responsible for a significant proportion of the heat transport.

Held and Hou Model: Asymmetric to Equator



(from James 1994)

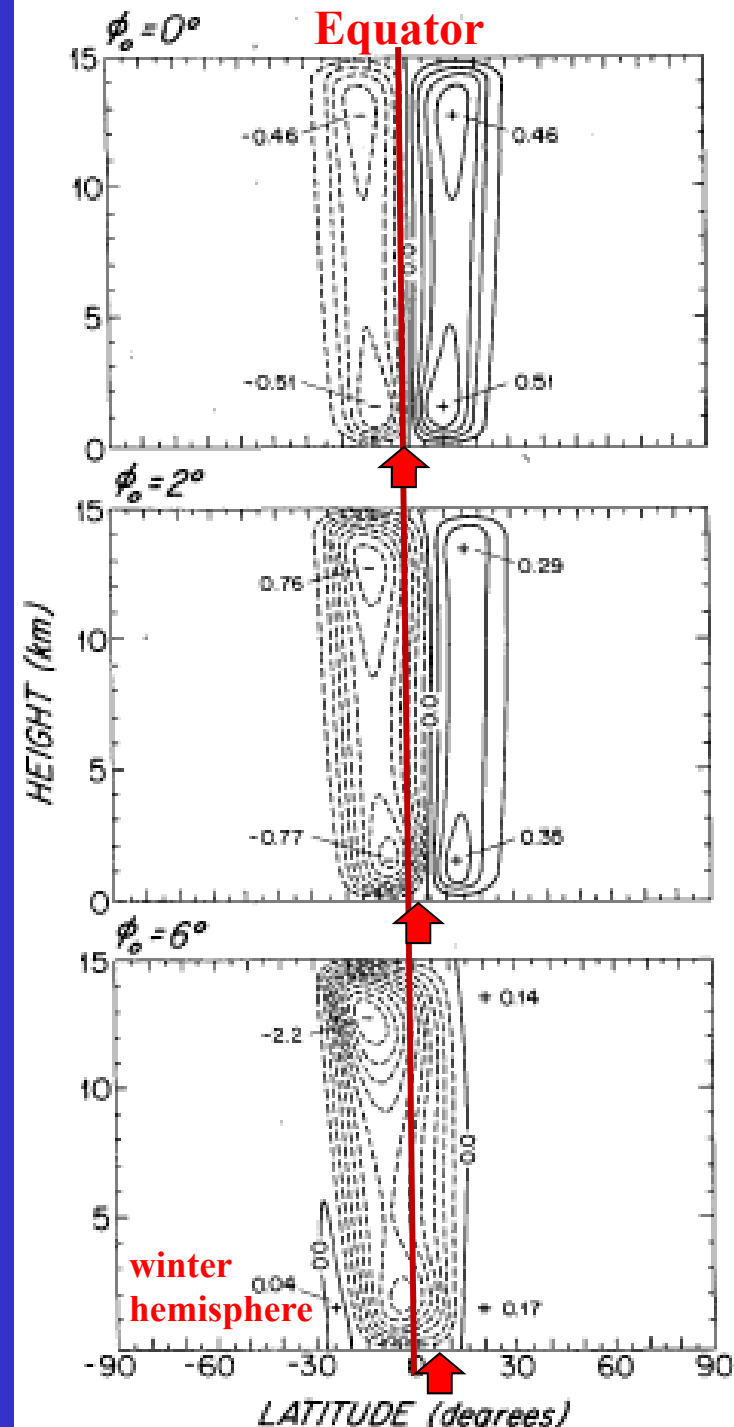
For maximum heating only 2-degree away from the equator, the winter cell is over three times as wide as the summer cell.



Off-Equatorial Heating

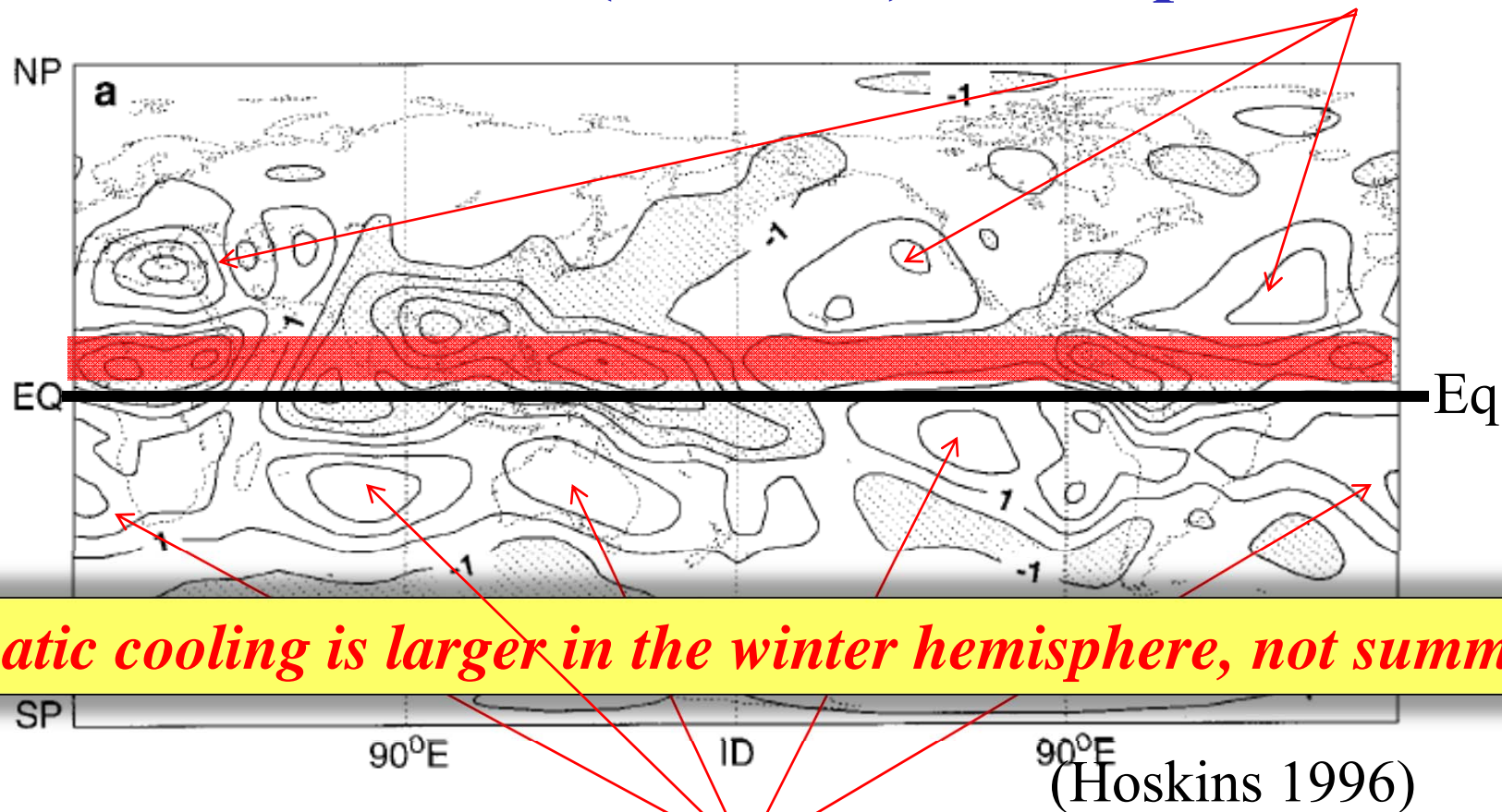
“ .. We find that moving peak heating even 2 degree off the equator leads to profound asymmetries in the Hadley circulation, *with the winter cell amplifying greatly and the summer cell becoming negligible.*”

--- Lindzen and Hou (1988; JAS)



Vertical Velocity $\omega_{500\text{mb}}$ (June-August 1994)

Northern (summer) subtropical descent

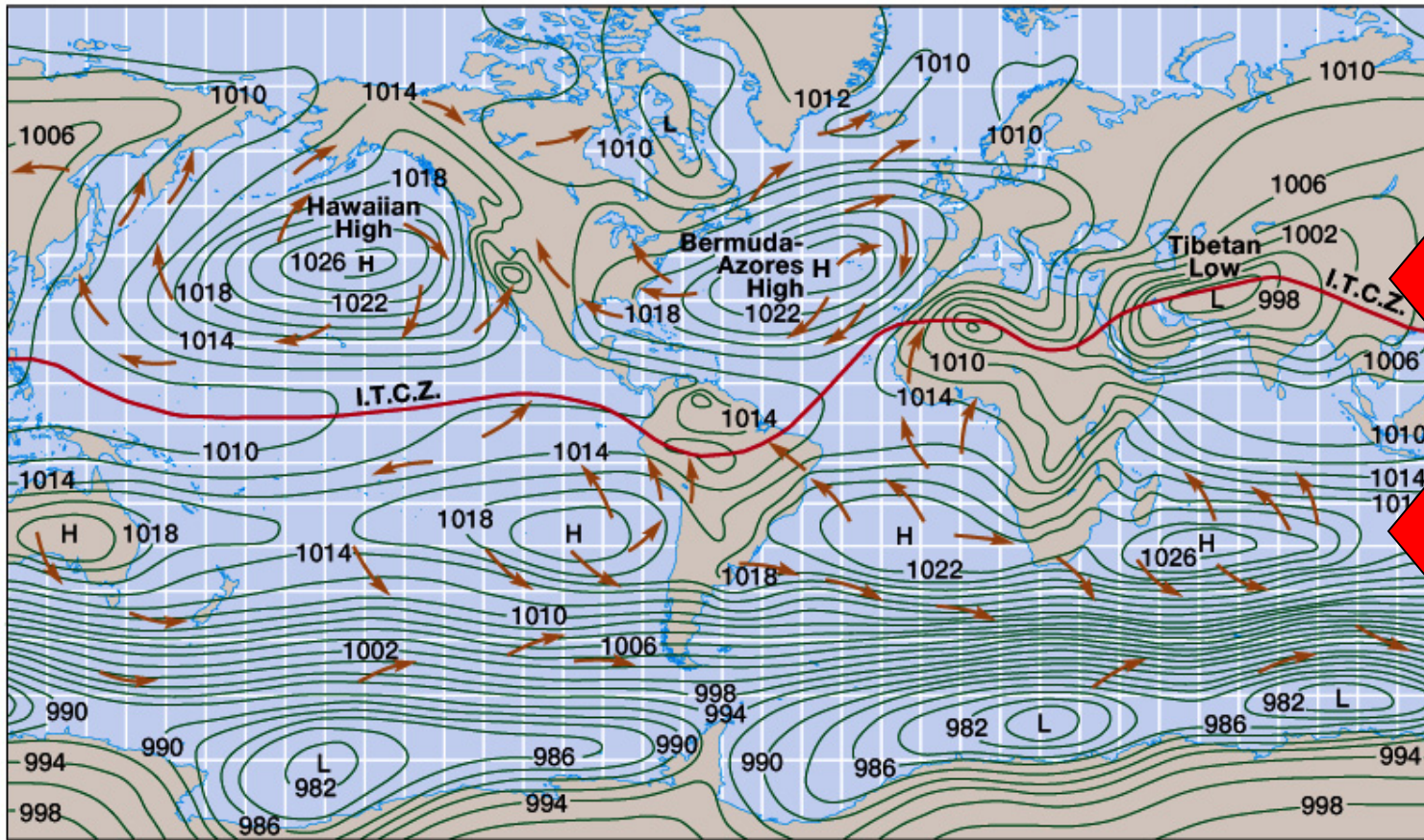


Diabatic cooling is larger in the winter hemisphere, not summer

Southern (winter) subtropical descent

Subtropical Highs

July (northern summer)



Localized Highs
(summer)

A Belt of Highs
(winter)

- *Winter subtropical highs can be explained by the Hadley circulation*
- *Summer subtropical highs has to be explained in the context of planetary waves*

Possible Mechanisms - Summer

The underlying mechanisms are still disputed:

- (1) Monsoon-desert mechanism (Rodwell and Hoskins 1996, 2001)
- (2) Local land-sea thermal contrast (Miyasaka and Nakamura 2005)
- (3) Diabatic amplification of cloud-reduced radiative cooling
- (4) Air-sea interaction

Monsoon-Desert Mechanism

(Rodwell and Hoskins 1996)

Asian monsoon

Desert/descending

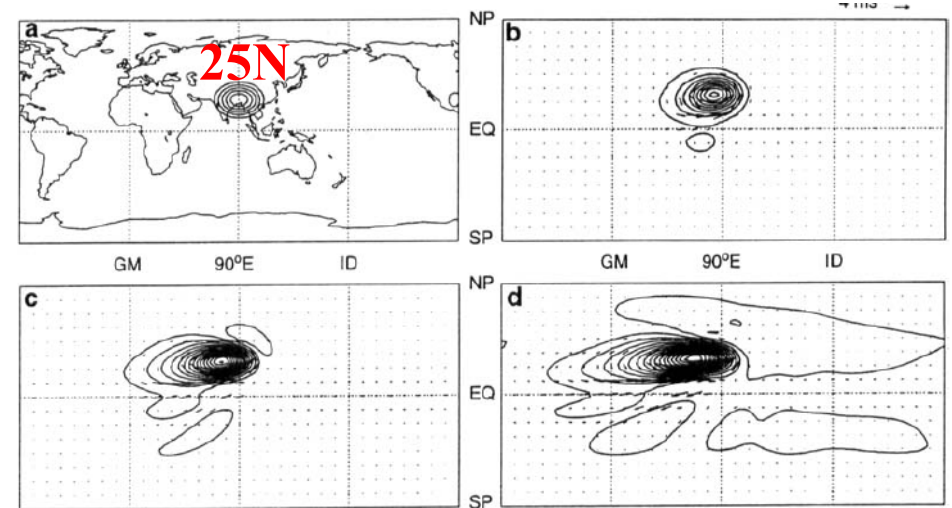
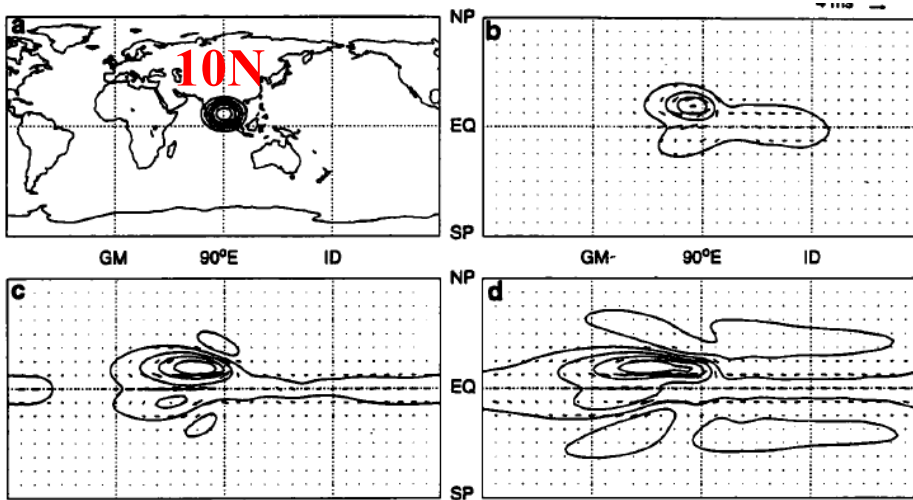
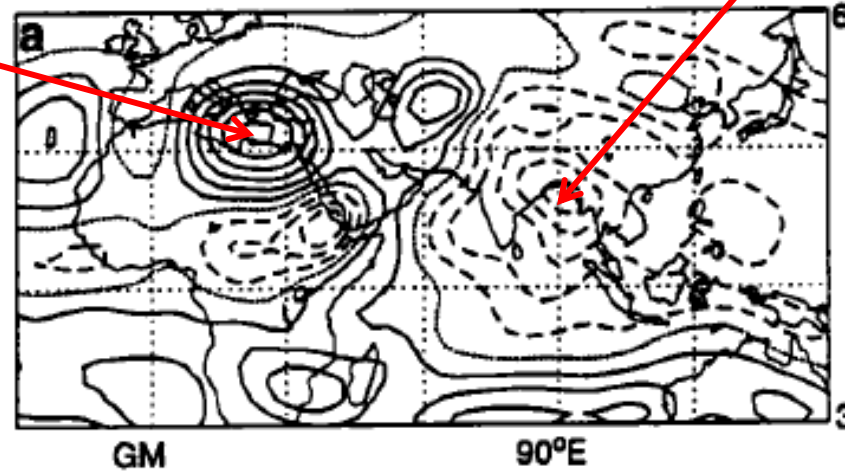
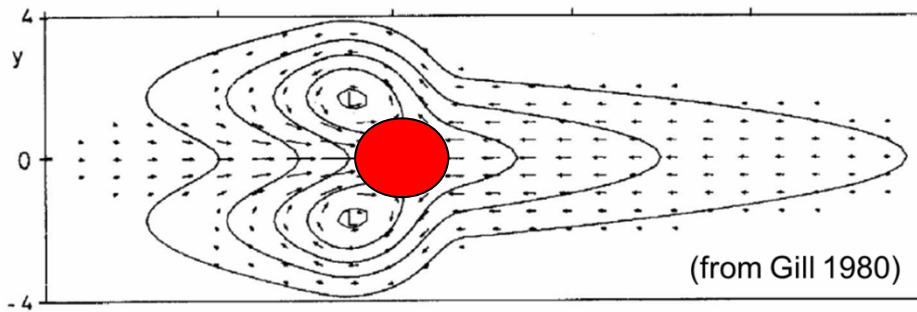


Figure 6. (a) Column-mean diabatic heating centred at 90°E , 10°N . The contour interval is 50 W m^{-2} ; the zero contour is not shown. (b), (c) and (d) show the corresponding perturbation surface pressure and 887 hPa horizontal winds for an integration linearized about a resting basic-state at (b) day 3, (c) day 7 and (d) day 11. The contour interval is 1 hPa .

Figure 7. (a) Column-mean diabatic heating centred at 90°E , 25°N . The contour interval is 50 W m^{-2} ; the zero contour is not shown. (b), (c) and (d) show the corresponding perturbation surface pressure and 887 hPa horizontal winds for an integration linearized about a resting basic-state at (b) day 3, (c) day 7 and (d) day 11. The contour interval is 1 hPa .